

SOME RIEMANNIAN MANIFOLDS ADMITTING A CONCIRCULAR SCALAR FIELD

MASAMI FUJII

Introduction

Recently, R. S. Kulkarni [2] and others have dealt with Riemannian manifolds admitting a concircular scalar field in a theory of curvature-preserving mappings or in connection with the so-called Nomizu's conjecture. Properties of concircular scalar fields, of special ones or in Einstein manifolds etc, were studied by Y. Tashiro [1] or other authors. However, we have less knowledge of concircular scalar fields in manifolds of constant scalar curvature, even in case of low-dimensional manifolds.

In this paper, we shall discuss properties of Riemannian manifolds admitting a concircular scalar field by using adapted coordinate systems for concircular scalar fields as tools. We refer to [1] as to notations and terminologies.

In 1, formulas with respect to an adapted coordinate system will be stated as preliminaries. In 2, we shall determine the structure of a 4-dimensional Riemannian manifold of constant scalar curvature admitting a concircular scalar field. We shall prove, in 3, that a 4-dimensional Einstein manifold admitting a concircular scalar field is of constant curvature, and slightly generalize Kulkarni's theorem concerning conformal map. It will be proved, in 4, that a manifold having the vanishing tensor $H_{\sigma\omega\nu\mu\lambda}$ defined by E. Cartan and admitting a concircular scalar field is of constant curvature, and, in 5, that a manifold satisfying $\overset{*}{H}_{\sigma\omega\mu\lambda} = H_{\sigma\omega\nu\mu\lambda}{}^\nu = 0$ is an Einstein manifold, unless the gradient vector field of the concircular scalar field is concurrent.

I would like to thank Professor Y. Tashiro who gave me continuous encouragements and valuable suggestions.

1. Preliminaries

We shall assume, throughout this paper, that a Riemannian manifold M is connected, differentiable and of dimension n , and the metric tensor $g_{\mu\lambda}$ of M is positive definite. Two kinds of indices run on the ranges

$$\begin{aligned} X, \lambda, \mu, \nu, \omega, \sigma, \tau &= 1, 2, 3, \dots, n \\ h, i, j, k &= 2, 3, \dots, n \end{aligned}$$

respectively. We denote the Christoffel symbol, the curvature tensor, the Ricci tensor and the scalar curvature of M by $\{\overset{\epsilon}{\mu\lambda}\}$, $K_{\nu\mu\lambda}{}^{\epsilon}$, $K_{\mu\lambda}$ and κ respectively, where the scalar curvature defined by $\kappa = \frac{1}{n(n-1)}K_{\mu\lambda}g^{\mu\lambda}$.

The scalar field ρ is said to be *concircular* if it satisfies the equation

$$(1.1) \quad \nabla_{\mu}\nabla_{\lambda}\rho = \phi g_{\mu\lambda},$$

∇ indicating covariant differentiation and ϕ being a scalar field, and to be *special concircular* if it satisfies the equation

$$(1.2) \quad \nabla_{\mu}\nabla_{\lambda}\rho = (-k\rho + b)g_{\mu\lambda}$$

k and b being constant.

Along any geodesic with arc-length u , the equation (1.1) becomes to the ordinary equation

$$(1.3) \quad \frac{d^2\rho}{du^2} = \phi.$$

We put $\rho_{,\lambda} = \delta_{,\lambda}\rho$. A point P is said to be *ordinary* or *stationary* according as $\rho_{,\lambda}(P) \neq 0$ or $\rho_{,\lambda}(P) = 0$. Stationary points of a concircular scalar field ρ are isolated and there exist at most two in M , see [1]. In a neighborhood U of an ordinary point of ρ , we can choose an adapted coordinate system (u^{ϵ}) having the following properties: the first coordinate u^1 is the arc-length u of ρ -curves, trajectories of ρ^{ϵ} , the coordinate hypersurfaces $u^1 = \text{constant}$ are ρ -hypersurfaces defined by $\rho = \text{constant}$, the field ρ is a function of $u^1 = u$ only, and the metric form ds^2 of M is given in the form

$$(1.4) \quad ds^2 = du^2 + \rho'(u)^2 \overline{ds}^2,$$

where the prime indicates ordinary derivative with respect to u and $\overline{ds}^2 = f_{ji} du^j du^i$ is a metric form of an $(n-1)$ -dimensional manifold \overline{M} . We indicate quantities of \overline{M} by barring. With respect to an adapted coordinate system (u^{ϵ}) , the metric tensor has components

$$(1.5) \quad g_{11} = 1, \quad g_{j1} = g_{1j} = 0, \quad g_{ji} = \rho'^2 f_{ji},$$

the curvature tensor $K_{\nu\mu\lambda}{}^{\epsilon}$ of M has components

$$(1.6) \quad \begin{aligned} K_{1j1}{}^h &= -K_{j11}{}^h = \frac{\rho'''}{\rho'} \delta_j^h, & K_{1ji}{}^h &= -\overline{K}_{j1i}{}^h = -\rho' \rho'' f_{ji}, \\ K_{kji}{}^h &= \overline{K}_{kji}{}^h - \rho''^2 (\delta_k^h f_{ji} - \delta_j^h f_{ki}), \end{aligned}$$

the other components being zero, the Ricci tensor has components

$$(1.7) \quad \begin{aligned} K_{11} &= -(n-1)\frac{\rho'''}{\rho'}, & K_{j1} &= K_{1j} = 0 \\ K_{ji} &= \bar{K}_{ji} - [(n-2)\rho''^2 + \rho'\rho''']f_{ji}, \end{aligned}$$

and the scalar curvature κ of M is equal to

$$(1.8) \quad \kappa = \frac{1}{n(\rho')^2} [(n-2)(\kappa - \rho''^2) - 2\rho'\rho'''],$$

where $\bar{\kappa}$ is the scalar curvature of \bar{M} defined by $\bar{\kappa} = \frac{1}{(n-1)(n-2)} \bar{K}_{ji} f^{ji}$.

2. 4-dimensional Riemannian manifolds of constant scalar curvature

Let M be a 4-dimensional Riemannian manifold of constant scalar curvature κ and ρ a conrcular scalar field. For $n=4$, the equation (1.8) is reduced to

$$(2.1) \quad 2\rho\rho'' + \frac{1}{2}(\rho''^2)' = \bar{\kappa}.$$

Since the left hand side depends on u only and $\bar{\kappa}$ is independent of u , $\bar{\kappa}$ is also a constant. According the signature of the constant scalar curvature κ , we put

$$(2.2) \quad \kappa = \begin{cases} \text{(I)} & 0 \\ \text{(II)} & -c^2 \\ \text{(III)} & c^2 \end{cases}$$

c being a positive constant. By a suitable choice of the arclength u , the general solution of (2.1) is given by one of

$$(2.3) \quad \rho'' = \begin{cases} \text{(I, A)} & au & (\bar{\kappa} = 0) \\ \text{(I, B)} & \bar{\kappa}u^2 + a & (\bar{\kappa} \neq 0) \\ \text{(II, A}_0) & a \exp 2cu - \frac{\bar{\kappa}}{2c^2} \\ \text{(II, A}_-) & a \sinh 2cu - \frac{\bar{\kappa}}{2c^2} \\ \text{(II, B)} & a \cosh 2cu - \frac{\bar{\kappa}}{2c^2} \\ \text{(III)} & a \cos 2cu + \frac{\bar{\kappa}}{2c^2}. \end{cases}$$

Therefore the manifold M has a local structure such that the metric form is given by (1. 4) substituted with (2. 3) for ρ'^2 .

Next we suppose that M is complete. Then the arc-length of any geodesic is extendable to the infinities. Since ρ -curves are geodesic, the cases (I, A) and (II, A₀) do not occur, and in the other cases, the inequalities

$$(2. 4) \quad \begin{cases} \text{(I, B)} & \bar{\kappa} > 0, & a \geq 0 \\ \text{(II, A}_0\text{)} & a > 0, & \bar{\kappa} < 0 \\ \text{(II, B)} & a > 0, & \bar{\kappa} \leq 2ac^2 \\ \text{(III)} & \bar{\kappa} > 0, & \bar{\kappa} \geq 2ac^2 \end{cases}$$

should be satisfied respectively, because $\rho'^2 \geq 0$.

Moreover, in order that there exists no stationary point of ρ in a complete manifold M , it is necessary and sufficient that the equalities in (2. 4) do not appear in all cases. Then the manifold M is topologically the direct product $I \times \bar{M}$ of a straight line I and a 3-dimensional complete manifold \bar{M} . By transferring the factor $\bar{\kappa}$ in the case (I, B) or a in the cases (II, A₀), (II, B) and (III) into the metric tensor f_{ji} of \bar{M} , in other words, applying a homothety to \bar{M} , the metric form of M is given by

$$(2. 5) \quad ds^2 = \begin{cases} \text{(I, B)} & du^2 + (u^2 + a)\bar{d}s^2 & (a > 0) \\ \text{(II, A}_0\text{)} & du^2 + \left(\exp 2cu - \frac{\bar{\kappa}}{2c^2}\right)\bar{d}s^2 & (\bar{\kappa} < 0) \\ \text{(II, B)} & du^2 + \frac{1}{2}\left(\cosh 2cu - \frac{\bar{\kappa}}{c^2}\right)\bar{d}s^2 & (\bar{\kappa} < c^2) \\ \text{(III)} & du^2 + \frac{1}{2}\left(\frac{\bar{\kappa}}{c^2} - \cos 2cu\right)\bar{d}s^2 & (\bar{\kappa} > c^2) \end{cases}$$

in the whole manifold M , respectively. On the other hand, the existence of a stationary point of ρ is possible in the cases (I, B), (II, B) and (III). Then M is of constant curvature and the scalar curvature is equal to $\bar{\kappa} = 1$ in (I, B), or $\bar{\kappa} = c^2$ in (II, B) and (III). There is one stationary point corresponding to $u = 0$ in (I, B) and (II, B) and are two corresponding to $u = 0$ and $u = \frac{\pi}{c}$ in (III). The metric form of M is given by

$$(2. 6) \quad ds^2 = \begin{cases} \text{(I, B)} & du^2 + u^2\bar{d}s^2 \\ \text{(II, B)} & du^2 + (\sinh cu)^2\bar{d}s^2 \\ \text{(III)} & du^2 + (\sin cu)^2\bar{d}s^2. \end{cases}$$

These are the polar forms of the metrics of (I, B) a Euclidean space, (II, B) a hyperbolic space and (III) a sphere, respectively. Thus we have established the following

Theorem 1. *Let M be a 4-dimensional complete Riemannian manifold of constant scalar curvature κ and suppose that M admits a concircular scalar field ρ . If there exists no stationary point of ρ , then the manifold M is topologically the direct product of a straight line I and a 3-dimensional complete manifold \bar{M} of constant scalar curvature $\bar{\kappa}$ and, the metric form of M is given by one of (2.5). If there exists a stationary point of ρ , then the manifold M is a Euclidean space, a hyperbolic space or a sphere.*

3. 4-dimensional Einstein manifolds

Let M be an n -dimensional Einstein manifold admitting a concircular scalar field ρ . Applying Ricci's formula to the equation (1.1), we have

$$(3.1) \quad -K_{\nu\mu\lambda}{}^{\kappa}\rho_{\kappa} = \phi_{\nu}g_{\mu\lambda} - \phi_{\mu}g_{\nu\lambda},$$

and contracting with $g^{\mu\lambda}$,

$$(3.2) \quad -K_{\nu}{}^{\kappa}\rho_{\kappa} = (n-1)\phi_{\nu}.$$

Since M is an Einstein manifold, that is, $K_{\nu}{}^{\kappa} = (n-1)\kappa\phi_{\nu}{}^{\kappa}$, we have the equation

$$(3.3) \quad \phi_{\nu} = -\kappa\rho_{\nu}, \quad \text{or} \quad \phi = -\kappa\rho + b$$

and

$$(3.4) \quad \nabla_{\mu}\nabla_{\lambda}\rho = (-\kappa\rho + b)g_{\mu\lambda},$$

where b is an integral constant. Hence, in an Einstein manifold, a concircular field is special and the characteristic constant of ρ is equal to the constant scalar curvature κ .

Theorem 2. *If a 4-dimensional Einstein manifold M admits a concircular scalar field, then the manifold is of constant curvature.*

Proof. With respect to an adapted coordinate system (u^{κ}) in a neighborhood U of any ordinary points of M , the equation (3.4) is reduced to the ordinary equation

$$(3.5) \quad \rho'' = -\kappa\rho + b \quad \text{and} \quad \rho''' = -\kappa\rho'.$$

Substituting these into the third equation of (1.7) for $n=4$, we have

$$(3.6) \quad K_{ji} = \bar{K}_{ji} - (2\rho^{l''} + \rho^l \rho^{l''}) f_{ji} = \bar{K}_{ji} - (2\rho^{l''} - \kappa \rho^{l''}) f_{ji}.$$

Since M is an Einstein manifold, we substitute $K_{ji} = 3\kappa g_{ji} = 3\kappa \rho^{l''} f_{ji}$ into (3.6) and obtain the equations

$$(3.7) \quad \bar{K}_{ji} = (3\kappa \rho^{l''} + 2\rho^{l''} - \kappa \rho^{l''}) f_{ji} = 2(\kappa \rho^{l''} + \rho^{l''}) f_{ji}.$$

This shows that \bar{M} is a 3-dimensional Einstein manifold

$$(3.8) \quad \bar{K}_{ji} = 2 f_{ji}$$

and constant scalar curvature $\bar{\kappa}$ is equal to

$$(3.9) \quad \bar{\kappa} = \kappa \rho^{l''} + \rho^{l''}.$$

As it is known that a 3-dimensional Einstein manifold is of constant curvature, the manifold \bar{M} is of constant curvature, that is,

$$(3.10) \quad \bar{K}_{kji}{}^h = \bar{\kappa} (\delta_k^h f_{ji} - \delta_j^h f_{ki}).$$

Substituting this and (3.9) into the third equation of (1.6), we have

$$(3.11) \quad \begin{aligned} K_{kji}{}^h &= (\bar{\kappa} - \rho^{l''}) (\delta_k^h f_{ji} - \delta_j^h f_{ki}) = \kappa \rho^{l''} (\delta_k^h f_{ji} - \delta_j^h f_{ki}) \\ &= \kappa (\delta_k^h g_{ji} - \delta_j^h g_{ki}). \end{aligned}$$

The first and second equations of (1.6) are rewritten as

$$(3.12) \quad \begin{aligned} K_{1j1}{}^h &= -\delta_j^h = \kappa (\delta_1^h g_{j1} - \delta_j^h g_{11}) \\ K_{1j1}{}^h &= \kappa \rho^{l''} f_{j1} = \kappa (\delta_1^1 g_{j1} - \delta_j^1 g_{11}) \end{aligned}$$

by means of (1.5). The equations (3.11) and (3.12) together make the tensor equation $K_{\nu\mu\lambda}{}^{\kappa} = \kappa (\delta_{\nu}^{\kappa} g_{\mu\lambda} - \delta_{\mu}^{\kappa} g_{\nu\lambda})$. Therefore the manifold M is of constant curvature at ordinary points. Since the stationary point of ρ is isolated if there is any, M is of constant curvature.

Q. E. D.

By virtue of this theorem, we give a slight generalization of Kul-karni's theorem [2] in a different way.

Corollary. *Let M and M^* be 4-dimensional Einstein manifolds which are nowhere of constant curvature. Then every conformal map of M into M^* is a homothety.*

Proof. Let f be a conformal map of M into M^* , and denote the metric tensor f^*g by components $g_{\mu\lambda}^*$. Then they are related by the

equation $g_{\mu\lambda}^* = \frac{1}{\rho^2} g_{\mu\lambda}$, where ρ is a positive valued scalar field. We indicate by asterisking quantities of $g_{\mu\lambda}^*$ corresponding to those of $g_{\mu\lambda}$. Then we obtain the transformation formulas

$$(3.13) \quad \left\{ \begin{matrix} \kappa \\ \mu\lambda \end{matrix} \right\}^* = \left\{ \begin{matrix} \kappa \\ \mu\lambda \end{matrix} \right\} - \frac{1}{\rho} (\delta_{\mu}^{\kappa} \rho_{\lambda} + \delta_{\lambda}^{\kappa} \rho_{\mu} - g_{\mu\lambda} \rho^{\kappa}),$$

$$(3.14) \quad K_{\nu\mu\lambda}^* = K_{\nu\mu\lambda} + \frac{1}{\rho} (\delta_{\nu}^{\kappa} \nabla_{\mu} \rho_{\lambda} - \delta_{\mu}^{\kappa} \nabla_{\nu} \rho_{\lambda} + g_{\mu\lambda} \nabla_{\nu} \rho^{\kappa} - g_{\nu\lambda} \nabla_{\mu} \rho^{\kappa}) - \frac{1}{\rho^2} \rho_{\omega} \rho^{\omega} (\delta_{\nu}^{\kappa} g_{\mu\lambda} - \delta_{\mu}^{\kappa} g_{\nu\lambda}),$$

$$(3.15) \quad K_{\mu\lambda}^* = K_{\mu\lambda} + \frac{1}{\rho} 2 \nabla_{\mu} \rho_{\lambda} + \frac{1}{\rho} g_{\mu\lambda} \nabla_{\kappa} \rho^{\kappa} - \frac{1}{\rho^2} 3 \rho_{\kappa} \rho^{\kappa} g_{\mu\lambda}.$$

Since M and M^* are Einstein manifolds, we substitute $K_{\kappa\lambda}^* = 3\kappa^* g_{\mu\lambda}^* = 3\kappa^* \rho^{-2} g_{\mu\lambda}$ and $K_{\mu\lambda} = 3\kappa g_{\mu\lambda}$ into (3.15), and obtain the following equation

$$(3.16) \quad \nabla_{\kappa} \rho_{\lambda} = \frac{\rho}{2} (3\rho^{-2} \kappa^* - 3\kappa - \frac{1}{\rho} \nabla_{\kappa} \rho^{\kappa} + \frac{3}{\rho^2} \rho^{\kappa}) g_{\mu\lambda}.$$

This equation means that ρ is a concircular scalar field if ρ would not be constant, and M would be a manifold of constant curvature by Theorem 2. This is a contradiction. Therefore ρ must be a constant, that is, f is a homothety.

Q. E. D.

We notice that κ^* need not be equal to κ .

4. Manifolds of $H_{\sigma\omega\nu\mu\lambda}^{\kappa} = 0$

E. Cartan defined the tensor $H_{\sigma\omega\nu\mu\lambda}^{\kappa}$ by the equation

$$(4.1) \quad H_{\sigma\omega\nu\mu\lambda}^{\kappa} = K_{\sigma\omega\nu}^{\tau} K_{\tau\mu\lambda}^{\kappa} + K_{\sigma\omega\mu}^{\tau} K_{\tau\nu\lambda}^{\kappa} + K_{\sigma\omega\lambda}^{\tau} K_{\tau\nu\mu}^{\kappa} - K_{\sigma\omega\tau}^{\kappa} K_{\nu\mu\lambda}^{\tau}.$$

Theorem 3. *If M is a manifold with the property $H_{\sigma\omega\nu\mu\lambda}^{\kappa} = 0$ and admits a concircular scalar field ρ such that ϕ is not identically constant, then the manifold is of constant curvature.*

Proof. We refer to an adapted coordinate system (u^{κ}) in a neighborhood of any ordinary points of ρ and put the indices $\sigma = 1, \omega = k, \kappa = h, \lambda = i, \mu = j, \nu = 1$ in the equation (4.1). Taking account of the components (1.6) of the curvature tensor, we have

$$\frac{\rho'''}{\rho'} \{ \bar{K}_{kji}{}^h - \rho'' (\delta_k^h f_{ji} - \delta_j^h f_{ki}) \} - \frac{\rho'''}{\rho'} (\rho' \rho''') (\delta_j^h f_{ki} - \delta_k^h f_{ji}) = 0$$

or

$$(4.2) \quad \frac{\rho'''}{\rho'} \{ \bar{K}_{kji}{}^h - (\rho'' - \rho' \rho''') (\delta_k^h f_{ji} - \delta_j^h f_{ki}) \} = 0.$$

As $\rho''' \neq 0$, we have from (4.2) the equation

$$(4.3) \quad \bar{K}_{kji}{}^h = (\rho'' - \rho' \rho''') (\delta_k^h f_{ji} - \delta_j^h f_{ki}),$$

from which follows $\rho'' - \rho' \rho''' = \bar{\kappa}$. This implies that \bar{M} is a manifold of constant curvature. Substituting (4.3) into the third of (1.6), we have

$$K_{kji}{}^h = -\frac{\rho'''}{\rho'} (\delta_k^h g_{ji} - \delta_j^h g_{ki}).$$

From the first and second equations of (1.6) and the other components being zero, we can obtain the tensor equation

$$K_{\omega\mu\lambda}{}^\kappa = -\frac{\rho'''}{\rho'} (\delta_\nu^\kappa g_{\mu\lambda} - \delta_\mu^\kappa g_{\nu\lambda}).$$

Since a stationary point is isolated, M is a manifold of constant curvature.

Q. E. D.

We put $J = \{u \mid \rho'''(u) = 0\}$. If J contains intervals, we have $\rho = Au^2 + Bu + C$, where A, B and C are constant. At the points of the complement J^c of J , the equation $\rho'' - \rho' \rho''' = \bar{\kappa}$ is satisfied under initial conditions $\rho''(0) = 2A$, $\rho'(0) = B$ and $\rho(0) = C$ by a suitable choice of arc-length u . Then the solution is given by

$$\rho(u) = \frac{B^2}{(4A^2 - \bar{\kappa})} \left(2A \cosh \frac{\sqrt{4A^2 - \bar{\kappa}}}{B} u + \sqrt{4A^2 - \bar{\kappa}} \sinh \frac{\sqrt{4A^2 - \bar{\kappa}}}{B} u - 2A \right) + C.$$

So the differentiability is broken at the point of $u=0$. Therefore, J is equal to the whole straight line I or discrete. When J is the straight line, $\rho'''(u) = 0$ for every point of I , that is, $\rho'' = b$ and $\nabla_\mu \rho_\lambda = b g_{\mu\lambda}$, b being constant. It follows that ρ_λ is concurrent or parallel. If J is discrete, M is of constant curvature at any point.

5. Manifolds of $\bar{H}_{\sigma\omega\mu\lambda}^* = 0$

We put the tensor,

$$(5.1) \quad H_{\sigma\omega\mu\lambda} = H_{\sigma\omega\nu\mu\lambda}{}^\nu = K_{\sigma\omega\mu}{}^\nu K_{\nu\lambda} + K_{\sigma\omega\lambda}{}^\nu K_{\mu\nu}.$$

Theorem 4. *If M is a manifold having the property $\overset{*}{H}_{\sigma\omega\mu\lambda}=0$ and admits a concircular scalar field ρ such that ϕ is not identically constant, then M is an Einstein manifold.*

Proof. Referring to an adapted coordinate system (u^a) in a neighborhood of any ordinary point of ρ , and putting the indices $\sigma=1$, $\omega=j$, $\lambda=i$, $\mu=1$ in the equation (5.1), we have

$$(5.2) \quad \frac{\rho'''}{\rho'} \{ \bar{K}_{j1} - (n-2)(\rho''^2 - \rho'\rho''') f_{j1} \} = 0.$$

As $\rho''' \neq 0$, it follows from (5.2) that

$$(5.3) \quad \bar{K}_{j1} = (n-2)(\rho''^2 - \rho'\rho''') f_{j1}.$$

This implies that \bar{M} is an Einstein manifold and the scalar curvature is equal to $\bar{\kappa} = \rho''^2 - \rho'\rho'''$. Substituting (5.3) into the third of (1.7), we have

$$(5.4) \quad K_{j1} = -(n-1) \frac{\rho'''}{\rho'} g_{j1}.$$

From the first and second equations of (1.7), we have the tensor equation $K_{\mu\lambda} = -(n-1) \frac{\rho'''}{\rho'} g_{\mu\lambda}$ and hence M is an Einstein manifold.

Q. E. D.

When ρ''' vanishes, the same argument as that of 4 is applicable.

REFERENCES

- [1] Y. TASHIRO: Conformal transformations in complete Riemannian manifolds, the Study Group of Geometry, 1967.
- [2] R. S. KULKARNI: Curvature structures and conformal transformations, J. Differential Geometry, 4 (1967), 425-451.

TSUYAMA TECHNICAL COLLEGE

(Received April 18, 1972)